

# Oxygen Photolysis in Solid Molecular Hydrogen: Details Associated with the Production of Water

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# Outline

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- I. Introduction to cold chemistry and its significance
- II. A walk through the lab and the basics of spectroscopy and theory
- III. Discuss implications and interpretation of experimental results and our conclusions thus far

\*\*\*I want to teach you something about chemistry!\*\*\*

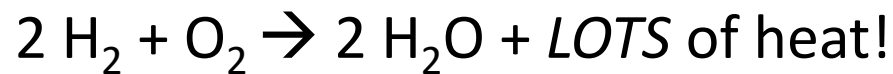
# Freezing an explosion - Impulse

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Endeavor shuttle  
STS-134 mission, May, 2011

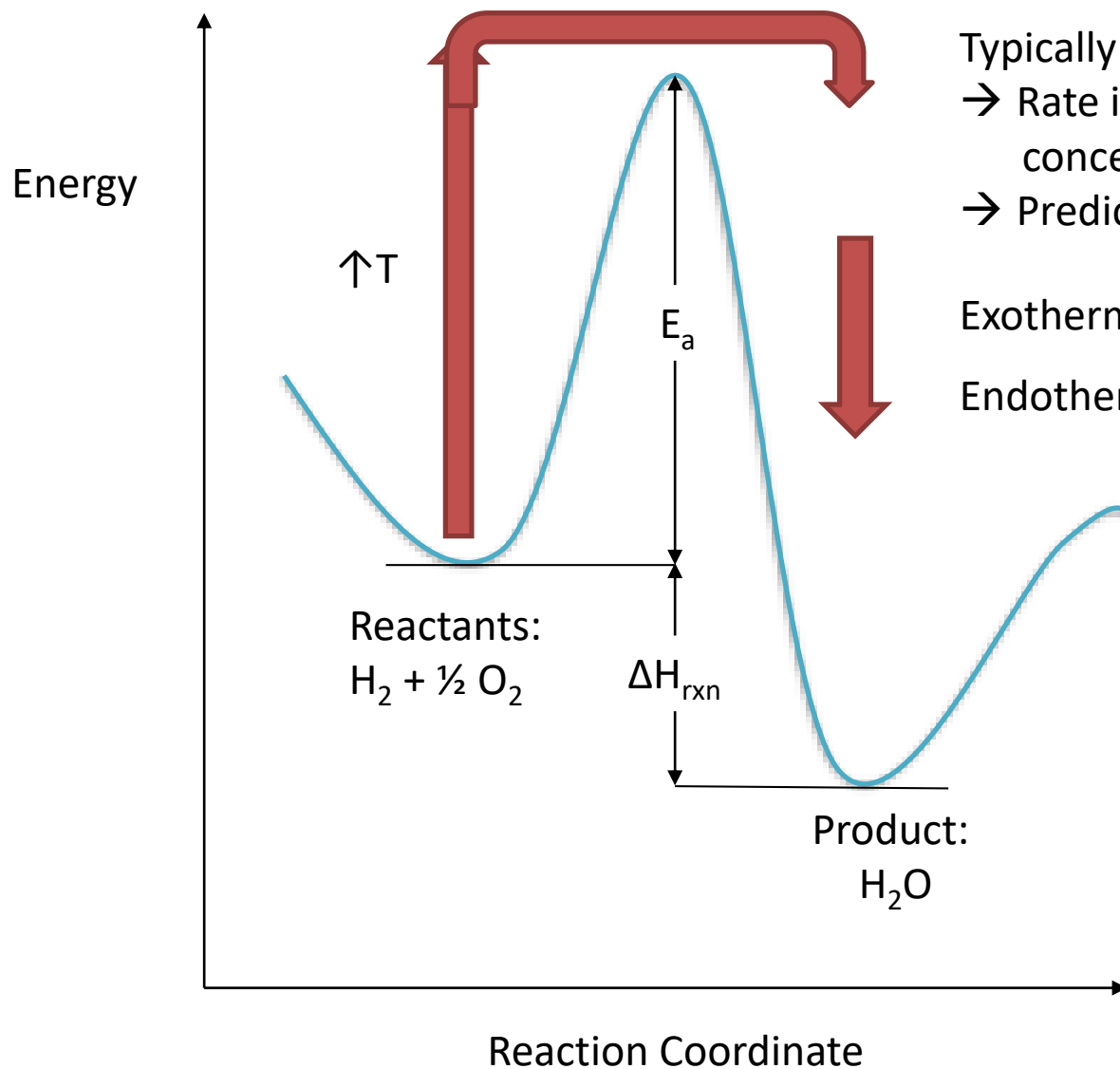
Consider the reaction:



- **Potential application: High energy density matter**

Ideal ratio: 1:1 O:H<sub>2</sub>

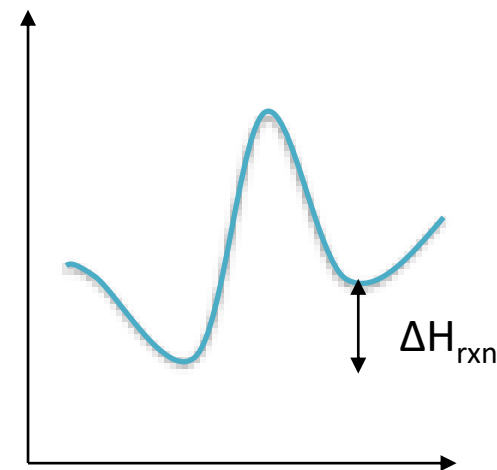
# Classical vs. Quantum Chemistry



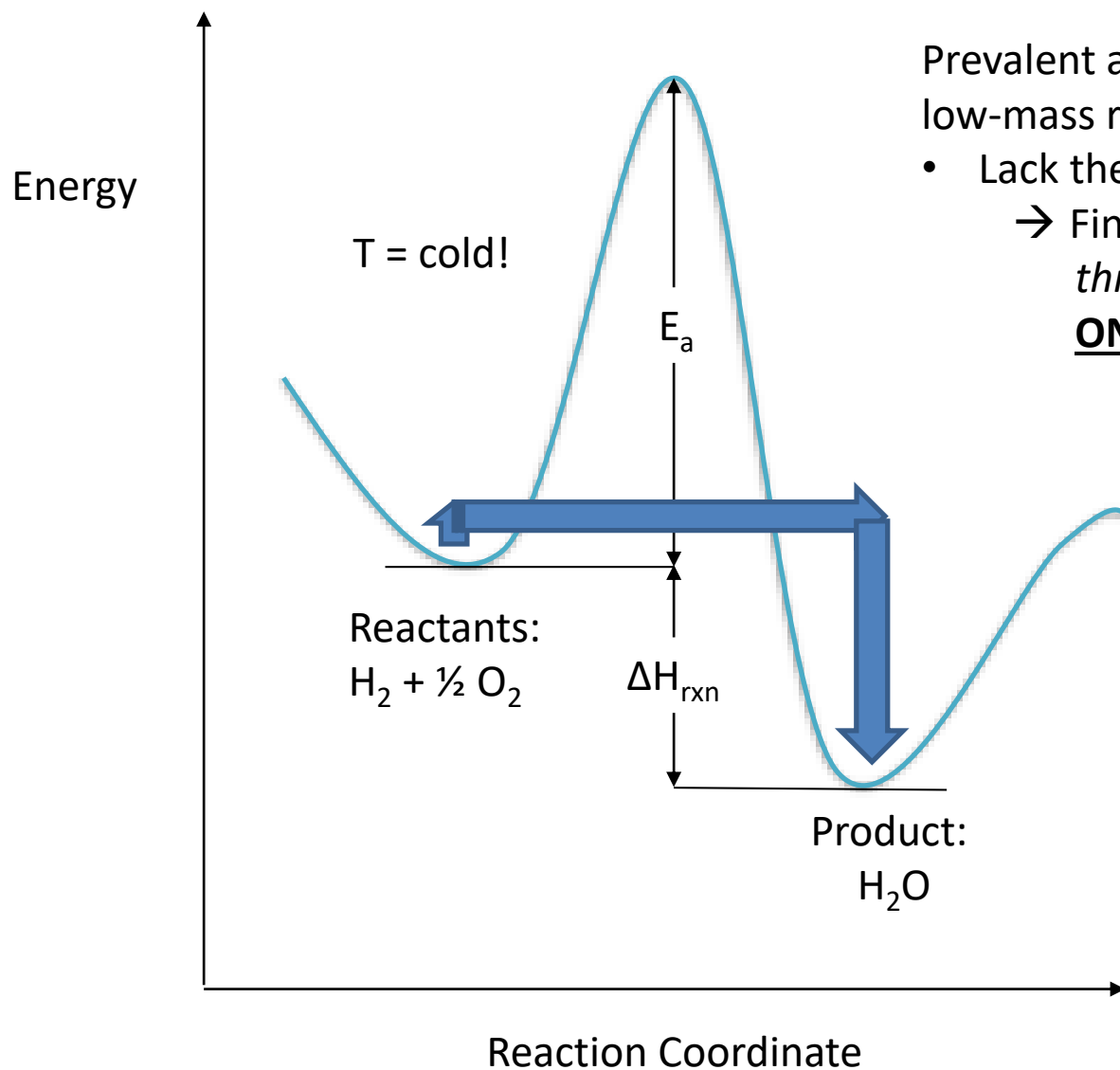
Typically add heat until reaction occurs  
→ Rate increases with temperature or concentration  
→ Predictable & well understood

Exothermic:  $0 > \Delta H_{rxn}$  (release energy)

Endothermic:  $0 < \Delta H_{rxn}$  (absorb energy)



# Classical vs. Quantum Chemistry



Prevalent at low temperatures and with low-mass reagents

- Lack thermal energy to surmount barrier  
→ Finite probability to tunnel *through* barrier...

**ONLY IF** reaction is exothermic!

# Interstellar Chemistry

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Photo credit: Peter Lik, "Stargazer"

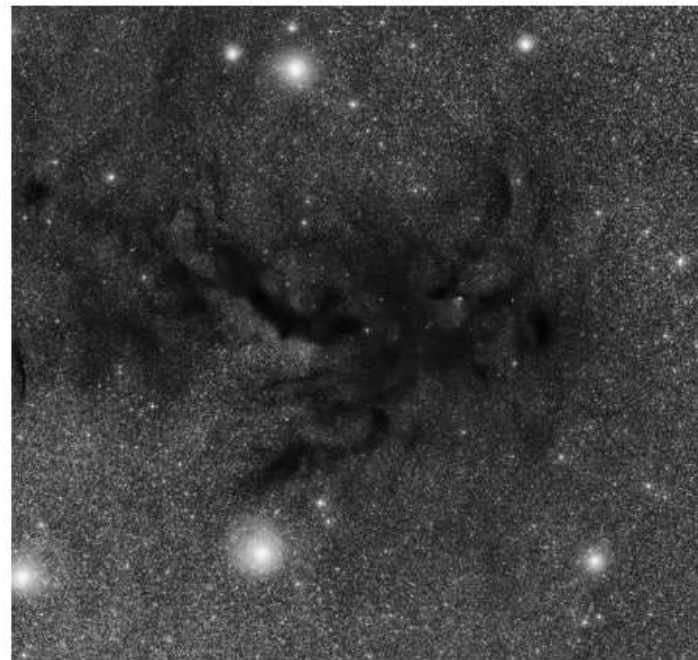
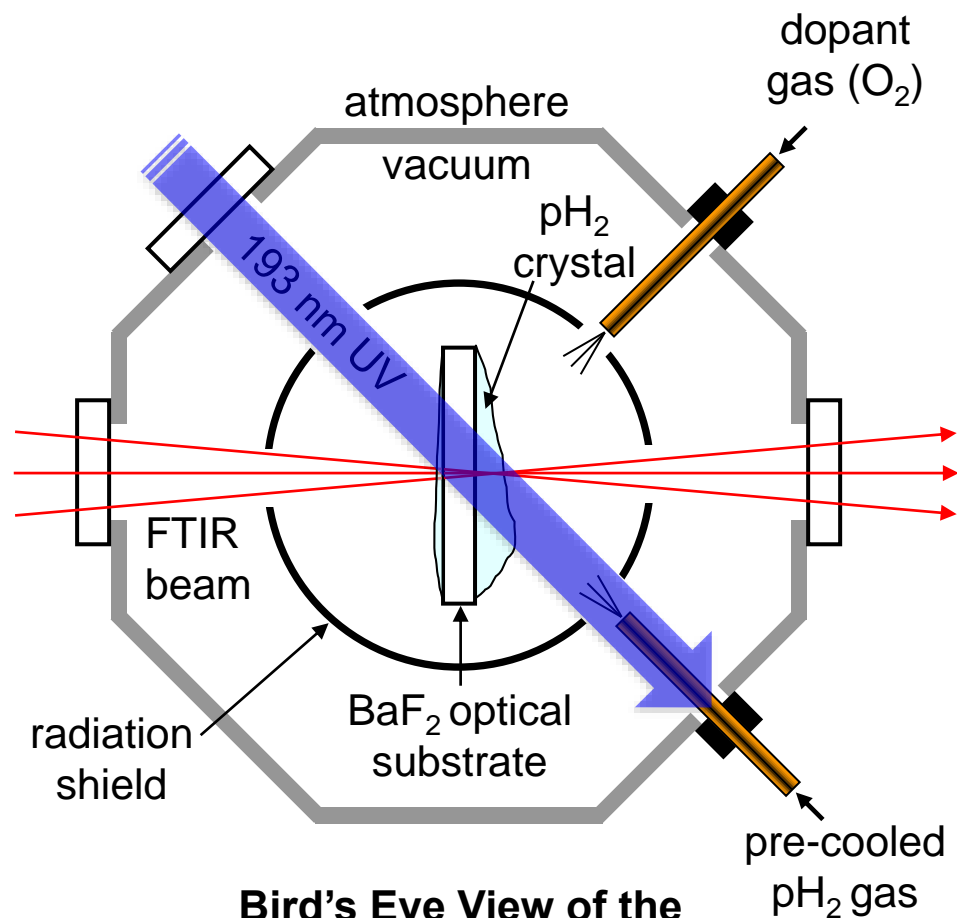
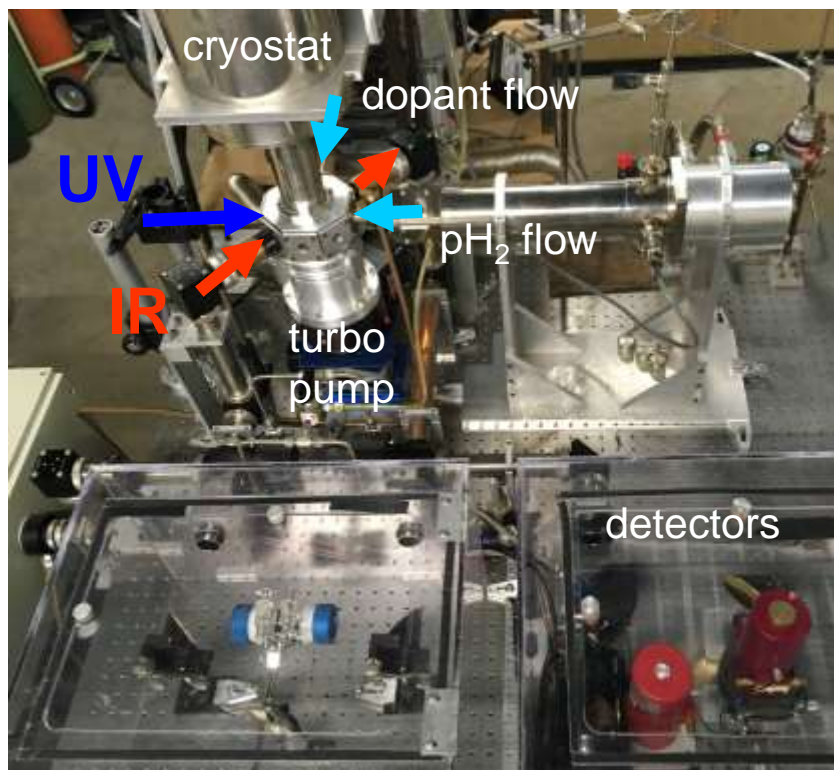


Fig. (3). This picture shows Barnard 59, part of a vast dark cloud of interstellar dust called the Pipe Nebula. The image was captured by the Wide Field Imager on the MPG/ESO 2.2-metre telescope at ESO's La Silla Observatory. Credit: ESO.

How does chemistry occur  
in interstellar space?

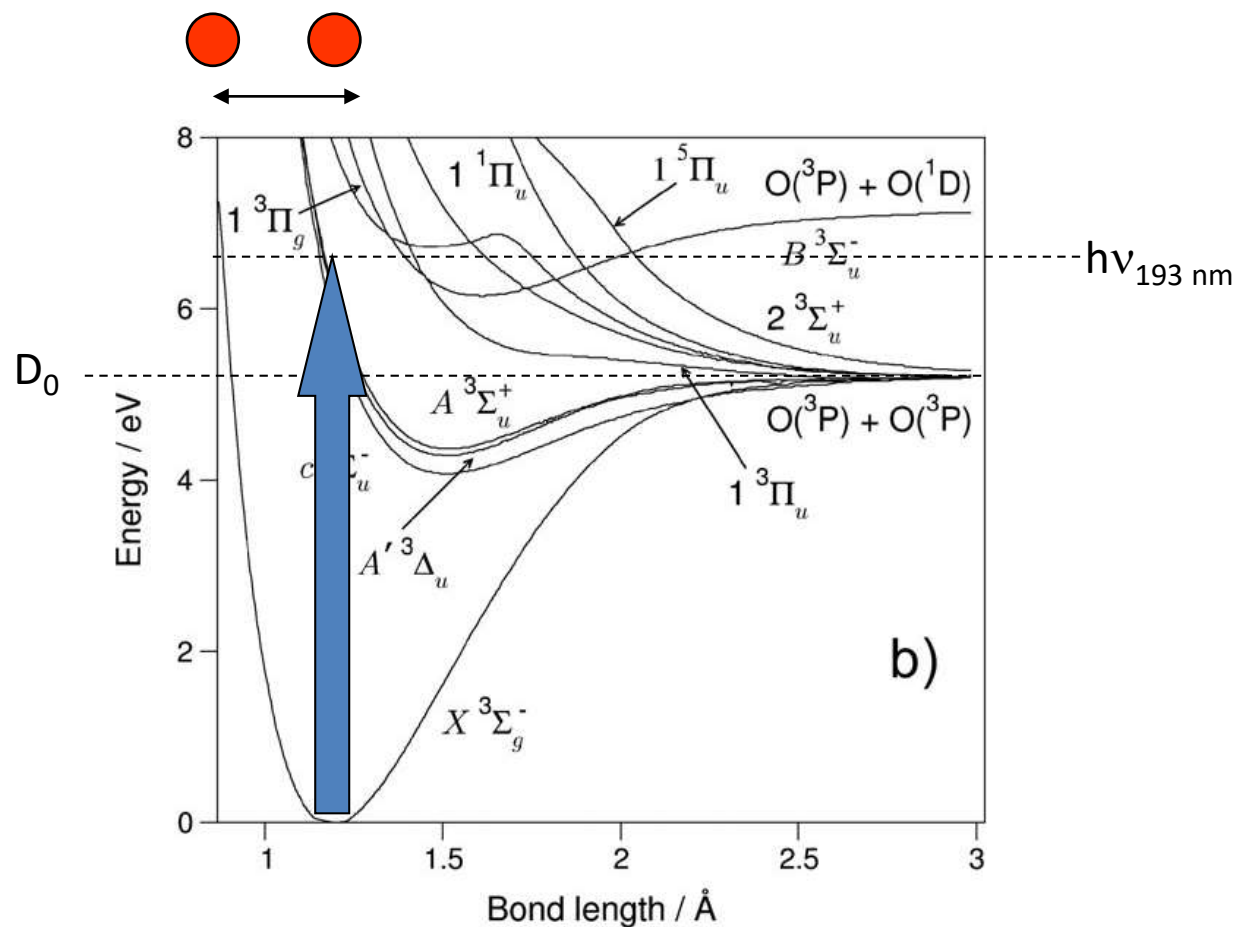
# Experimental Procedure (welcome to the lab!)



**Bird's Eye View of the  
Liquid Helium Bath Cryostat**  
(1.6 – 5 K; aka near -450 °F)

- Monitor chemical identity & concentration based on spectral frequency & intensity

# Photochemistry of O<sub>2</sub> inside solid hydrogen

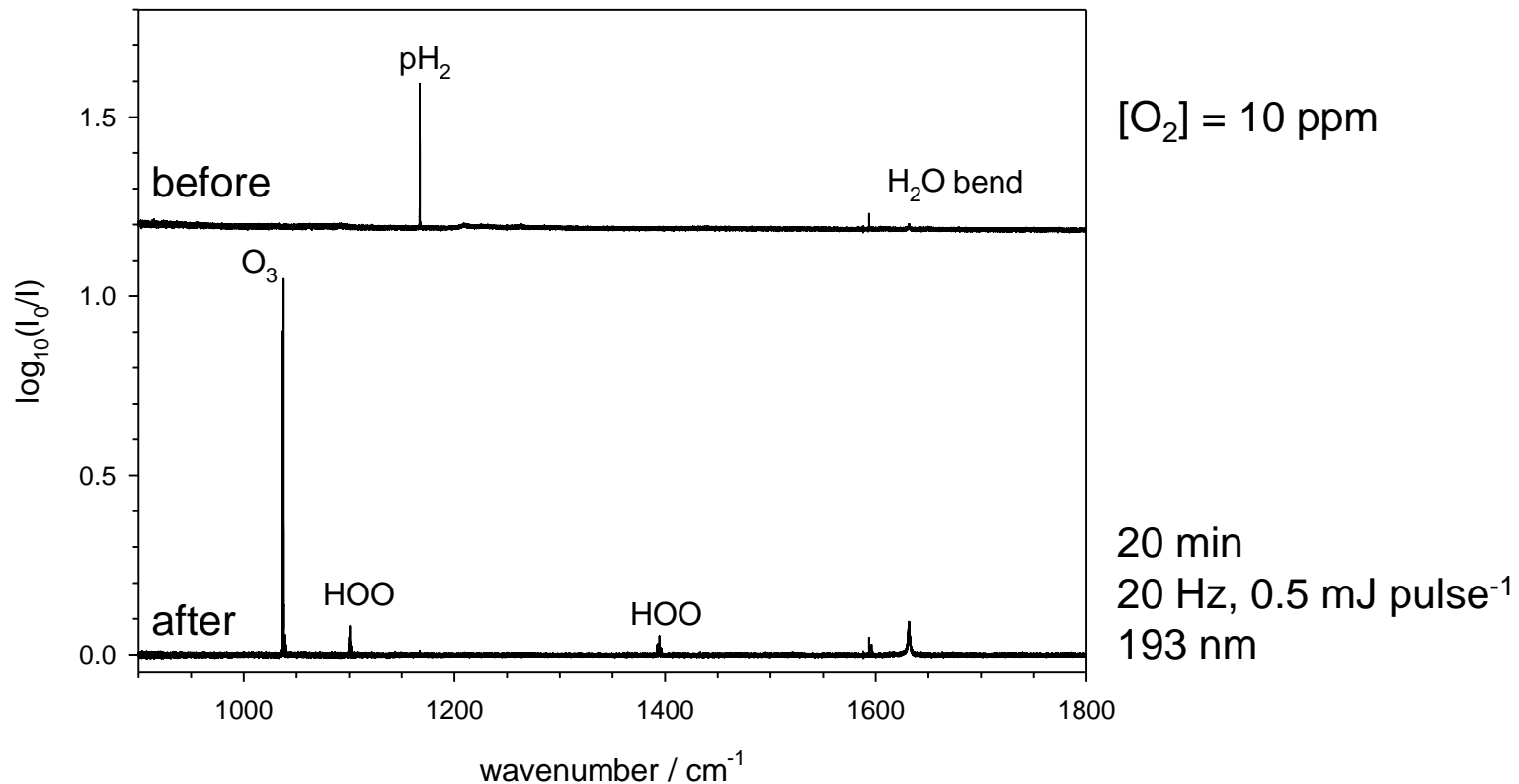
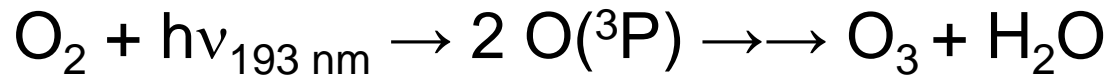


$$h\nu_{193 \text{ nm}} \approx 51,800 \text{ cm}^{-1}$$

$$D_0 = 5.12 \text{ eV} = 41,300 \text{ cm}^{-1}$$



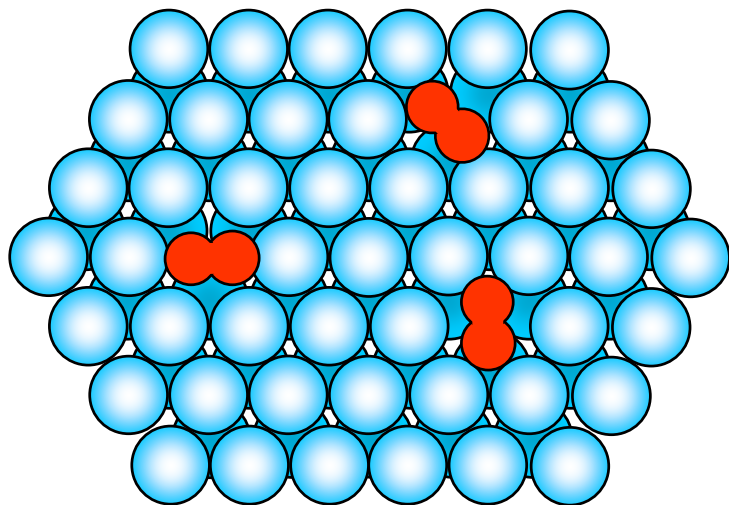
# Photochemistry of O<sub>2</sub> inside solid hydrogen



Assignments from O<sub>3</sub>/Ar studies: \*  $\nu_1 = 1105 \text{ cm}^{-1}$ ,  $\nu_3 = 1040 \text{ cm}^{-1}$

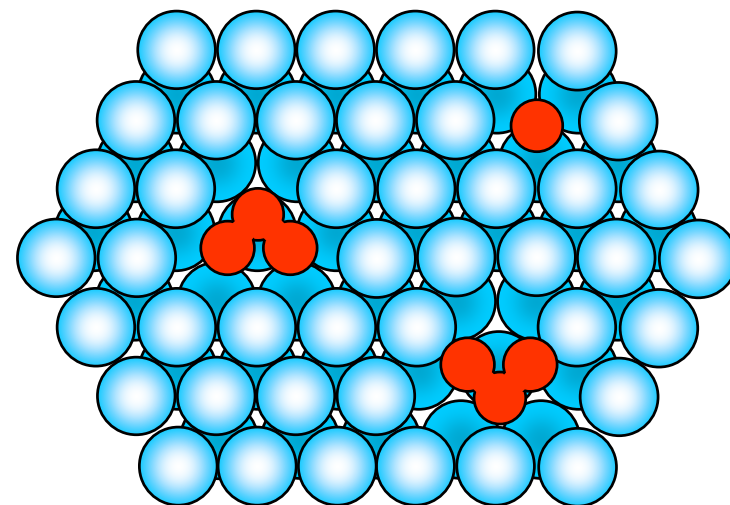
[\*] L. Andrews and R.C. Spiker, Jr., *JPC* **76**, 3208 (1972).

# Oxygen recombination dynamics in solid H<sub>2</sub>

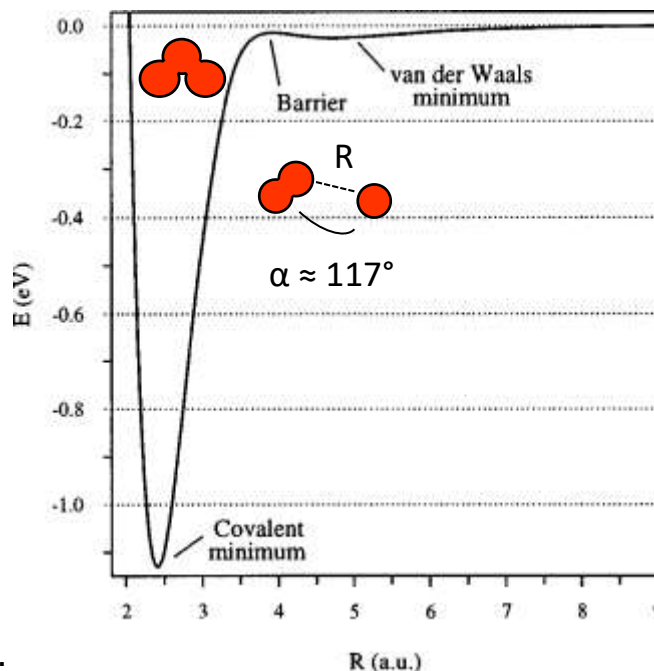
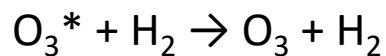
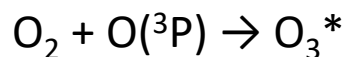


before

193 nm

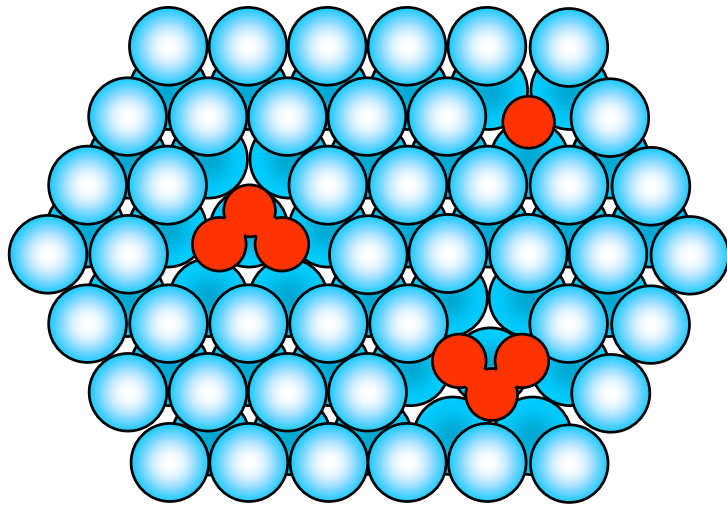


after



- Ozone formation provides indirect proof of O-atoms in solid pH<sub>2</sub>

# Formation of water in solid H<sub>2</sub>

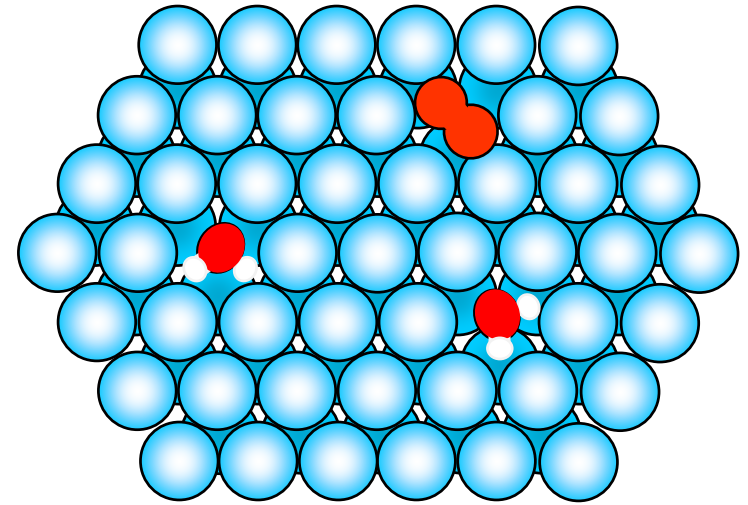


before

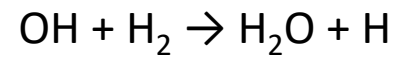
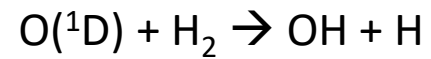
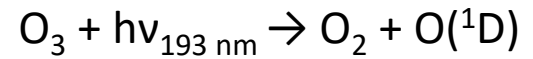
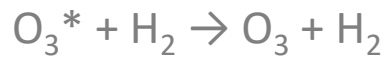
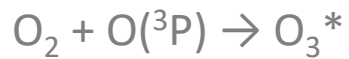
193 nm



(again\*)

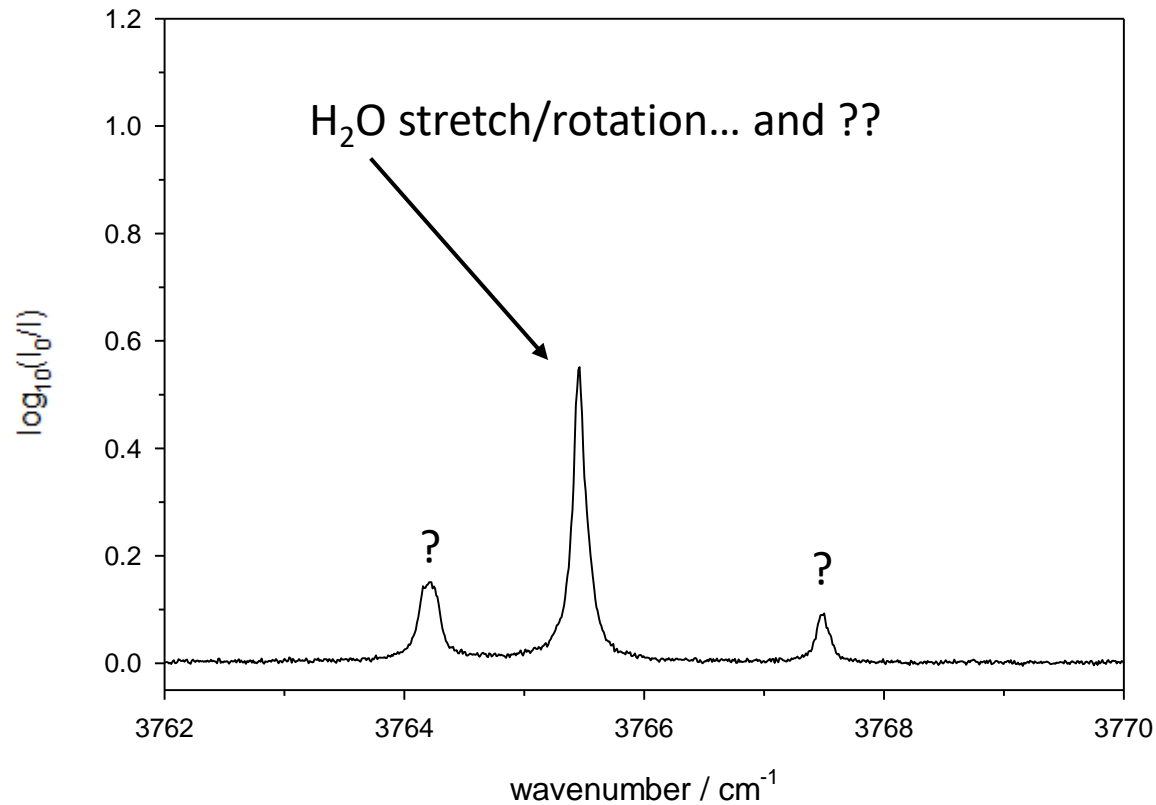


after



# Infrared satellite features of the water molecule

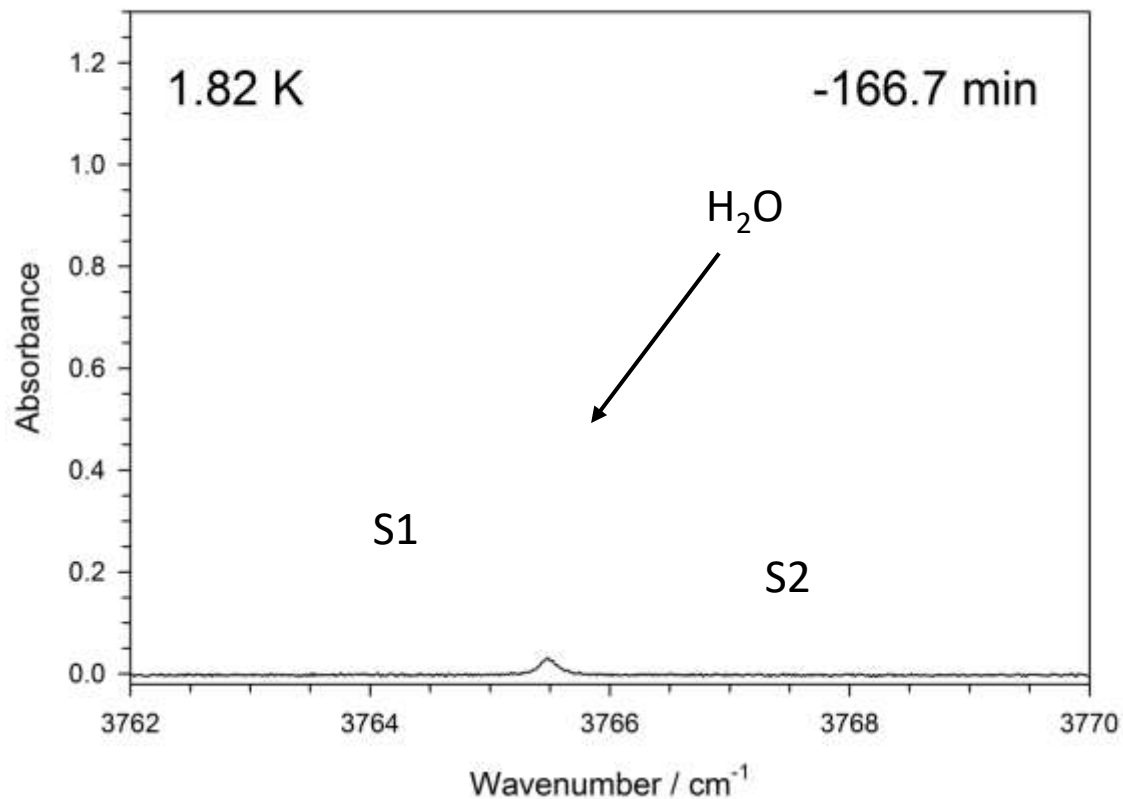
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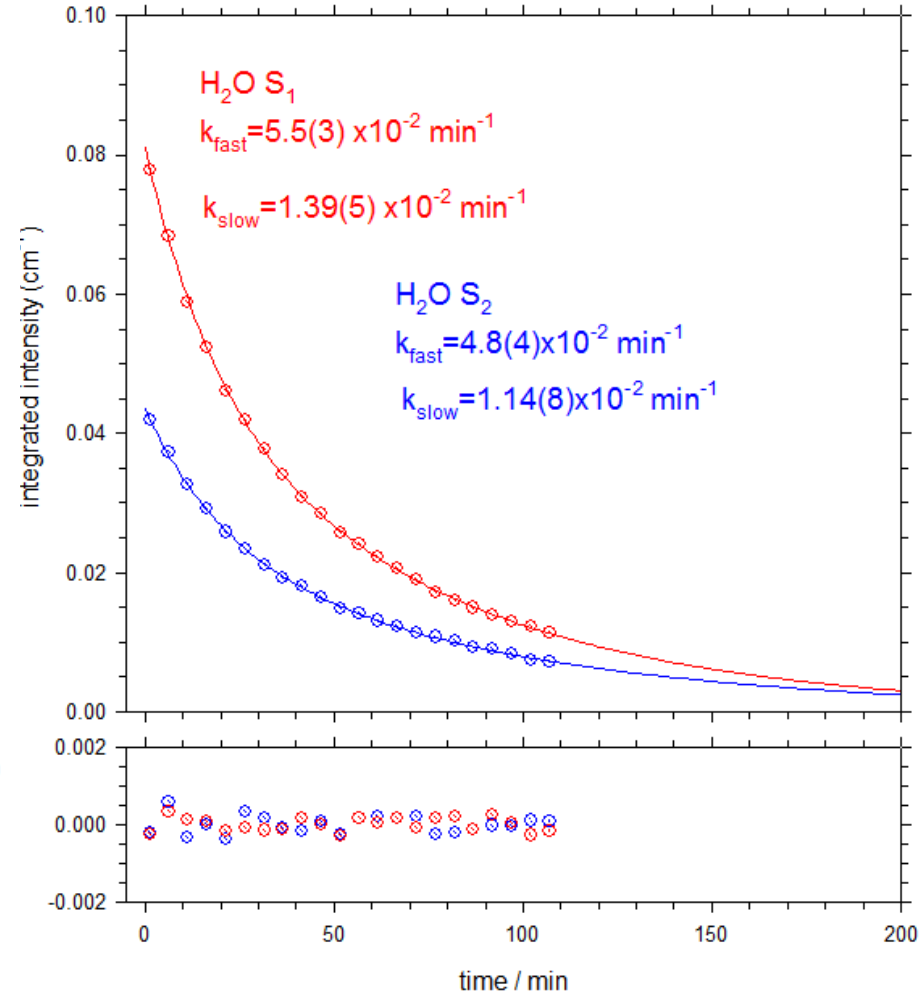
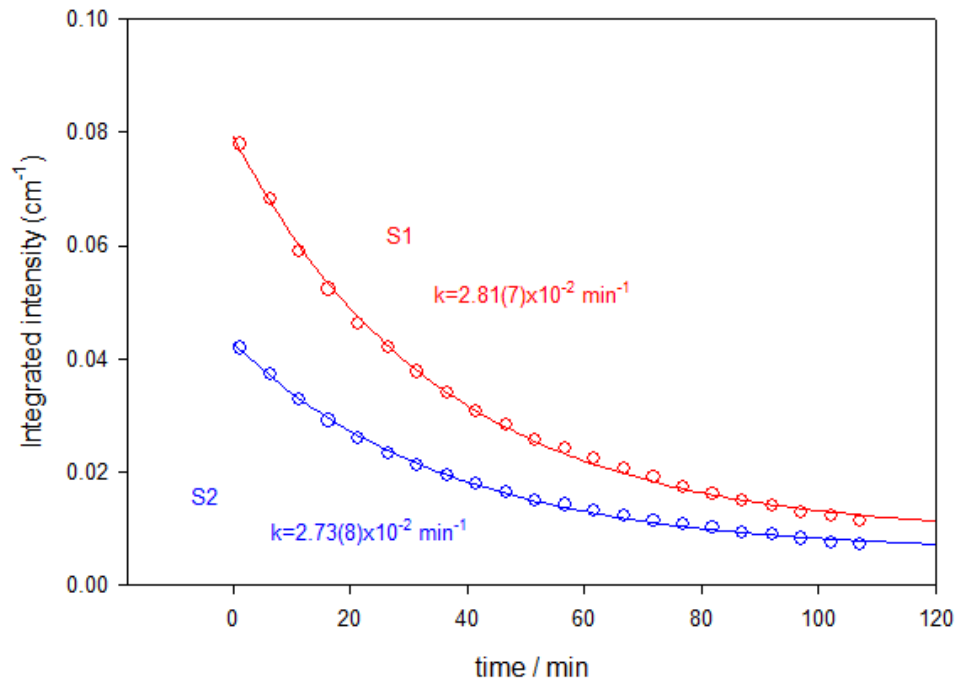
- not due to water/H<sub>2</sub> isomeric complexes

# Time-dependent changes after 193 nm UV photolysis

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# Tracking the decay of water satellite features



# Formation of water next to vacancies in solid H<sub>2</sub>

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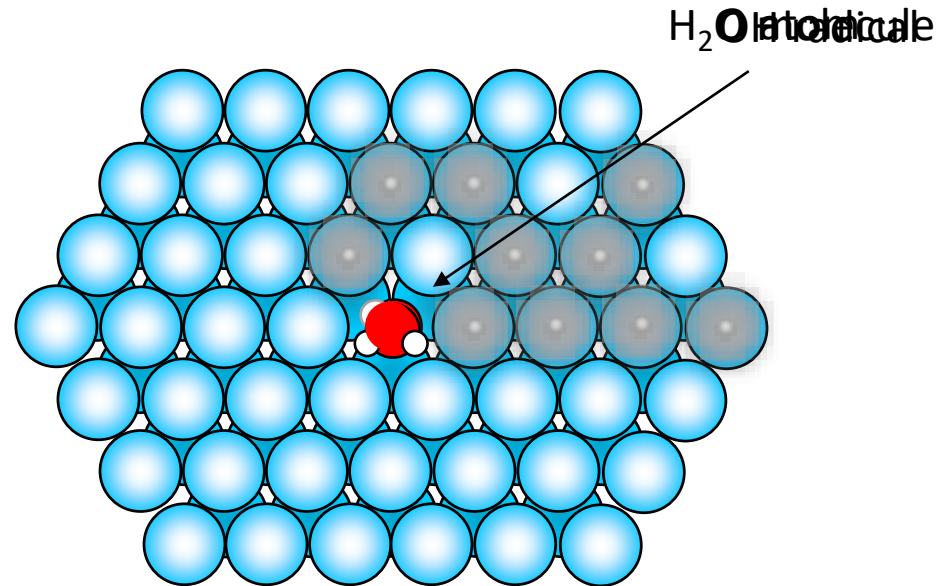
Find qualitative lifetime result:

$$\tau_{\text{water}} \approx 30\text{-}40 \text{ min}$$

Decay lasts longer in other systems:<sup>1</sup>

$$\tau_{\text{FA}} = 115.2 \pm 6.7 \text{ min}$$

Potential for unique chemical phenomena in various dopant-H<sub>2</sub>-host environments that produce water!<sup>2</sup>



Toy/PowerPoint  
model of the  
transformation

[1] D.T. Anderson *et. al.*, *JPC Lett.* **3**, 342-347 (2012).

[2] W.R. Wonderly and D.T. Anderson, *Low Temp. Phys.* **38**, 673-678 (2012).

# Conclusions

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- I. Form  $O_3$  from  $O_2$  photolysis - signaling production of O-atoms - and  $H_2O$  production follows photolysis of  $O_3$ 
  - Achieving a 1:1 O: $H_2$  concentration for use as rocket fuel is not feasible due to mobility & reactivity of O-atoms in cryogenic  $H_2$
- II. Recognize spectroscopic signatures of nascent  $H_2O$  within chemical environments that differ from equilibrium signals
- III. The time-dependent response of these water satellites could be anomalous when compared to previous findings – but, there is certainly evidence of low-temperature quantum-chemistry occurring within the hydrogen crystal



# Acknowledgements

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The Anderson Group - Fundamental Studies Near the Low Temperature Limit:



Left - Right: Myself, Morgan Balabanoff and Dr. Fred Mutunga



Dr. David Anderson

Financial support:

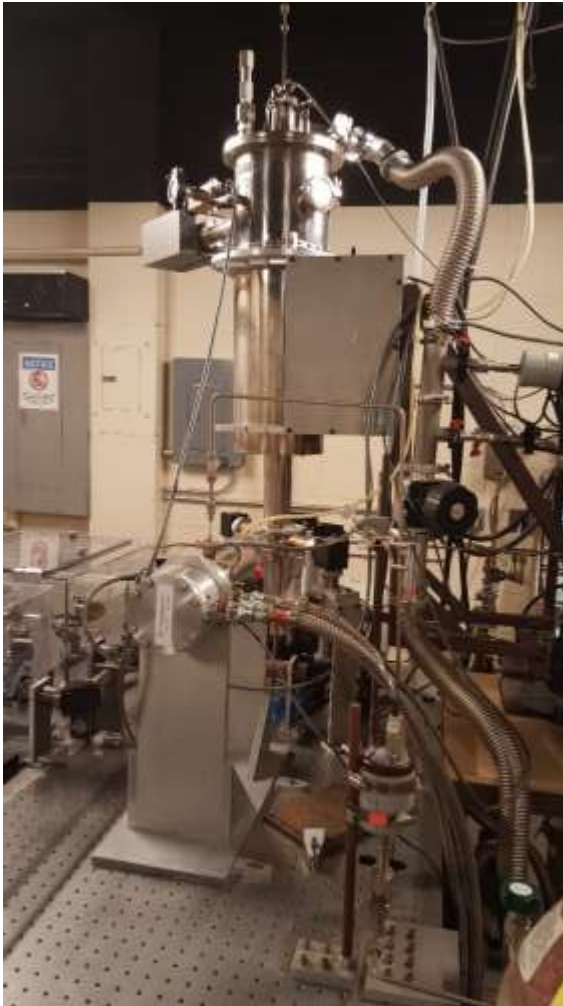


Questions or Comments?

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# Reaching Liquid Helium Temperatures

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Liquid Helium Bath Cryostat

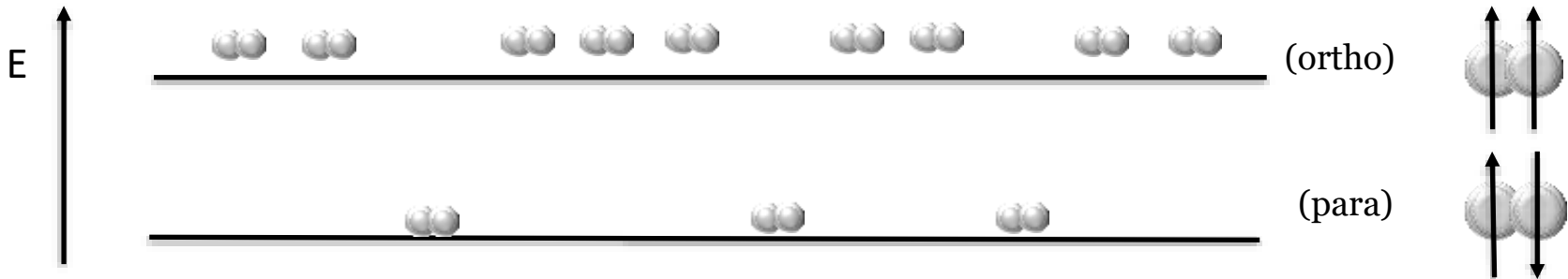
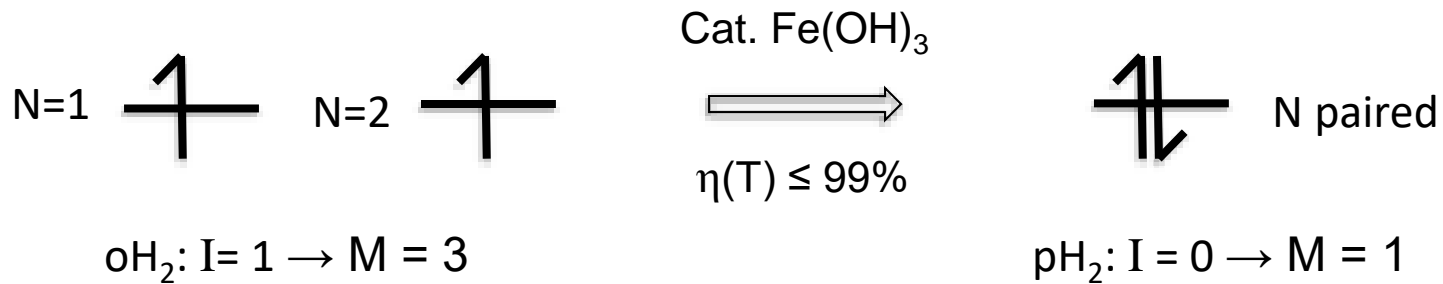
## Cool down procedure:

1. Precool with liquid N<sub>2</sub> to 77 K
2. Sap heat in cryostat with liquid He
  - He boiling point = 4.2 K (1 atm)
3. Fill bath and regulate pressure

Experimental range of 1.5 – 5.0 K.

# Ortho → Para Hydrogen Conversion

- “Normal” hydrogen consists of 3:1 ortho-para H<sub>2</sub>
- Consider the two nuclei (N) of H<sub>2</sub>
  - Multiplicity,  $M = 2I + 1$ ;  $I = \pm 1/2$  (Nuc. Spin Quantum #)
  - Fermions are antisymmetric w.r.t. exchange



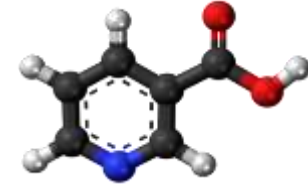
# Future Plans

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## Vitamin B<sub>3</sub> production in solid hydrogen

- An essential water soluble vitamin

Potential prebiotic seed/catalyst



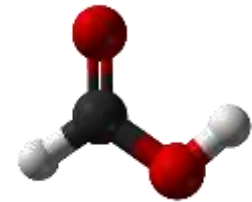
Vitamin B<sub>3</sub>

(a.k.a. niacin or nicotinic acid)

- 193 nm photolysis of pyridine (C<sub>5</sub>H<sub>5</sub>N) in pH<sub>2</sub>  
MeOH or FA / pyridine multi-dopant experiments

*e.g.*, FA + 193 nm → HCO, HOCO, H<sub>2</sub>O, etc.

HOCO + py. + 193 nm → Vit. B<sub>3</sub> (via *meta*-substitution)?



Formic  
Acid (FA)

Also extend to CO and O<sub>2</sub> non-polar ice mixtures

- Could this vitamin be formed via quantum tunneling reactions?